

Nuclear structure effects in light muonic atoms

Krzysztof Pachucki and Albert Wienczek

Faculty of Physics, University of Warsaw, Pasteura 5, 02-093 Warsaw, Poland

(Dated: April 28, 2015)

Nuclear structure corrections to energy levels of light muonic atoms are derived with particular attention to the nuclear mass dependence. The obtained result for the 2P-2S transition of 1.717(20) meV serves for determination of the nuclear charge radius from the spectroscopic measurement in muonic deuterium.

PACS numbers: 31.30.jr, 36.10.Ee 14.20.Dh

In order to resolve discrepancies for the proton charge radius [1–3], spectroscopic measurements in light muonic atoms, such as μD , $\mu^3\text{He}$, and $\mu^4\text{He}$ have been performed [4] for the comparison of nuclear charge radii with those obtained from traditional atomic spectroscopy or electron scattering from nuclei. The nuclear charge radius can be determined from spectroscopic measurements, provided the atomic structure is well known and the influence of nuclear excitation on atomic levels is properly accounted for. The atomic structure is well understood because one can calculate within quantum electrodynamics the atomic levels with very high precision, up to the value of fundamental constants. Much more problematic is the accurate description of nuclei and their electromagnetic interactions with surrounding electrons and muons, because of the difficulty in solving quantum chromodynamics in the low energy scale.

The nuclear polarizability effects in muonic atoms have been studied for some time. In 1977 Friar in [5] calculated the nonrelativistic nuclear electric dipole polarizability and Coulomb corrections for muonic helium. Eighteen years later, Leidemann and Rosenfelder in [6] calculated the inelastic contribution for μD in a more general approach by construction of the forward two-photon scattering amplitude for the deuteron. More recently, we calculated in [7] nuclear structure effects in muonic deuterium using a perturbative formalism and have shown the absence of the Zemach correction. The results of this perturbative approach have been confirmed by Friar in [8] using zero-range nucleon potentials. A systematic dispersion relation approach was used in [9] to obtain the complete two-photon exchange contribution, but the result suffered from insufficient inelastic scattering data from the deuteron. Recently, a perturbative approach has been pursued by independent derivation and numerical calculations for μHe in [10] and μD in [11].

In this work we include higher order terms in the expansion in a small parameter being the nuclear excitation energy over the muon mass, and recalculate all other contributions with special emphasis on the nuclear mass dependence and separation of the so-called pure recoil corrections. Since the nuclear effects are the main source of theoretical uncertainties in muonic atoms, we aim to calculate them as accurately as possible, in order to extract precise nuclear charge radii from the muonic atom spectroscopy. Our main limitation will come from the

simplified model of nuclear interaction with the electromagnetic field which assumes certain commutation relations, from the neglect of possible corrections to the electric dipole operator and from the uncertainty regarding the neutron polarizability.

In the following we derive general formulas for the nuclear polarizability shift using various perturbative expansions. We aim to improve results obtained in Refs. [7, 10, 11] by correcting mass dependencies and including higher order terms. Let us first introduce the notation used. Positions of the muon and nucleons are \vec{r}, \vec{r}_a . Corresponding relative positions with respect to the nuclear mass center are \vec{p}, \vec{p}_a . Momenta of the muon and nucleons are \vec{p}, \vec{p}_a . Relative nucleon momenta are $\vec{q}_a = \vec{p}_a - \vec{P} m_a / M$, where the total nuclear momentum is \vec{P} with the nuclear mass $M = \sum_a m_a$. The canonical commutation relations

$$[r_a^i, p_b^j] = i \delta_{ab} \delta^{ij} \quad (1)$$

for relative coordinates are the following:

$$[p_a^i, q_b^j] = i \left(\delta_{ab} - \frac{m_b}{M} \right) \delta^{ij}. \quad (2)$$

We assume that the nuclear Hamiltonian is of the form

$$\begin{aligned} \tilde{H}_N &= \sum_a \frac{\vec{p}_a^2}{2 m_a} + V_{\text{nucl}} \\ &= \frac{\vec{P}^2}{2 M} + \sum_a \frac{\vec{q}_a^2}{2 m_a} + V_{\text{nucl}} \\ &= \frac{\vec{P}^2}{2 M} + H_N, \end{aligned} \quad (3)$$

where m_a is a proton or a neutron mass. In what follows we will neglect the isospin number, so we will assume that each nucleon is a proton or a neutron. Under this assumption the electromagnetic interaction is local and is much easier to deal with. Later on, when matrix elements are calculated for the deuteron, the correct isospin number is assumed. This simplified treatment is because the full description of nuclear electromagnetic interactions, including separation of the center-of-mass motion [12], have not yet been presented in the literature.

We start derivation from the second-order Coulomb interaction in the nonrelativistic approximation

$$\delta E = \left\langle \phi \phi_N \left| \delta V \frac{1}{E_N + E_0 - H_N - H_0} \delta V \right| \phi \phi_N \right\rangle, \quad (4)$$

where

$$\delta V = \sum_{a=1}^Z \frac{\alpha}{|\vec{\rho} - \vec{\rho}_a|} - \frac{Z\alpha}{\rho}, \quad (5)$$

and where H_0 is the nonrelativistic Hamiltonian of the muon with the reduced mass

$$H_0 = \frac{\vec{p}^2}{2m_r} - \frac{Z\alpha}{\rho}. \quad (6)$$

The distance of protons from the nuclear mass center ρ_a is much smaller than that of the muon ρ , so the dominating contribution comes from the electric dipole excitations

$$\delta E = \alpha^2 \left\langle \phi \phi_N \left| \frac{\vec{d} \cdot \vec{\rho}}{\rho^3} \frac{1}{E_N + E_0 - H_N - H_0} \frac{\vec{d} \cdot \vec{\rho}}{\rho^3} \right| \phi \phi_N \right\rangle, \quad (7)$$

where $\vec{d} = \sum_{a=1}^Z \vec{\rho}_a$. Denoting the nuclear excitation energy by E , the nonrelativistic polarizability correction is

$$\delta E = \frac{\alpha^2}{3} \int_{E_T} dE |\langle \phi_N | \vec{d} | E \rangle|^2 \left\langle \phi \left| \frac{\vec{\rho}}{\rho^3} \frac{1}{E_0 - H_0 - E} \frac{\vec{\rho}}{\rho^3} \right| \phi \right\rangle \quad (8)$$

The nuclear excitation energy E is much larger than a typical muonic atomic excitation energy, thus one may perform the large E expansion of the muonic matrix element. The corresponding formula for this expansion is

$$\begin{aligned} \left\langle \phi \left| \frac{\vec{\rho}}{\rho^3} \frac{1}{H_0 - E_0 + E} \frac{\vec{\rho}}{\rho^3} \right| \phi \right\rangle &= 4\pi \phi^2(0) \sqrt{\frac{2m_r}{E}} \\ &+ c_1 \frac{(Z\alpha)^4 m_r^4}{E} + c_2 \frac{(Z\alpha)^5 m_r^4}{E} \sqrt{\frac{2m_r}{E}} + c_3 \frac{(Z\alpha)^6 m_r^5}{E^2}, \end{aligned} \quad (9)$$

where

$$\phi^2(0) = \frac{(m_r Z \alpha)^3}{\pi n^3} \delta_{l0}, \quad (10)$$

$$c_1(2P - 2S) = -\frac{1}{12} - \frac{1}{2} \ln\left(\frac{2m_r(Z\alpha)^2}{E}\right), \quad (11)$$

$$c_2(2P - 2S) = \frac{19}{32} + \frac{\pi^2}{12}, \quad (12)$$

$$c_3(2P - 2S) = -\frac{7}{6} + \frac{\zeta(3)}{2} + \frac{5}{8} \ln\left(\frac{2m_r(Z\alpha)^2}{E}\right) \quad (13)$$

From this expansion, the leading electric dipole polarizability contribution is [5]

$$\delta_0 E = -\frac{4\pi\alpha^2}{3} \phi^2(0) \int_{E_T} dE \sqrt{\frac{2m_r}{E}} |\langle \phi_N | \vec{d} | E \rangle|^2, \quad (14)$$

while Coulomb corrections are

$$\begin{aligned} \delta_{C1} E &= \frac{Z^4 \alpha^6 m_r^4}{6} \int_{E_T} \frac{dE}{E} \left[\frac{1}{6} + \ln\left(\frac{2m_r(Z\alpha)^2}{E}\right) \right] \\ &\quad \times |\langle \phi_N | \vec{d} | E \rangle|^2, \end{aligned} \quad (15)$$

$$\begin{aligned} \delta_{C2} E &= -\frac{Z^5 \alpha^7 m_r^3}{6} \left(\frac{19}{32} + \frac{\pi^2}{12} \right) \int_{E_T} dE \left(\frac{2m_r}{E} \right)^{3/2} \\ &\quad \times |\langle \phi_N | \vec{d} | E \rangle|^2. \end{aligned} \quad (16)$$

$\delta_{C3} E$ is small and thus can be neglected for light muonic atoms. The dipole operator \vec{d} in the above is the position of protons with respect to the nuclear mass center. However, one may expect some corrections to \vec{d} . Indeed the chiral effective field theory predicts various relativistic corrections to the electric dipole operator. We do not calculate them here and therefore associate a relative uncertainty of 1%, which is twice the binding energy per nuclear mass.

In the evaluation of further corrections we neglect Coulomb corrections, and so assume the on-mass-shell approximation for the muon. All corrections can therefore be expressed in terms of the two-photon forward scattering amplitude. Let us consider again the related muonic matrix element P for the nonrelativistic two-Coulomb exchange

$$P = \left\langle \phi \left| \frac{\alpha}{|\vec{\rho} - \vec{\rho}_a|} \frac{1}{(H_0 - E_0 + E)} \frac{\alpha}{|\vec{\rho} - \vec{\rho}_b|} \right| \phi \right\rangle. \quad (17)$$

Using the on-mass-shell approximation and subtracting the point Coulomb exchange it becomes

$$\begin{aligned} P &= \alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 \left(E + \frac{k^2}{2m_r} \right)^{-1} \\ &\quad \times (e^{i\vec{k} \cdot (\vec{\rho}_a - \vec{\rho}_b)} - 1). \end{aligned} \quad (18)$$

This integral can easily be performed analytically, but we will choose another way, which will be convenient when relativistic corrections are included. We will calculate directly the expansion coefficients in powers of E . There are two characteristic integration regions: $k \sim \sqrt{E}m$ and $k \sim m$, where m is the muon mass. In the first integration region, where k is small, one performs an exponent expansion in powers of $\vec{k} \cdot (\vec{\rho}_a - \vec{\rho}_b)$. The leading quadratic term is the electric dipole contribution

$$\begin{aligned} P_0 &= -\frac{4\pi}{3} \alpha^2 \phi^2(0) \sqrt{\frac{2m_r}{E}} \frac{(\vec{\rho}_a - \vec{\rho}_b)^2}{2} \\ &\rightarrow \frac{4\pi}{3} \alpha^2 \phi^2(0) \sqrt{\frac{2m_r}{E}} \vec{\rho}_a \cdot \vec{\rho}_b \end{aligned} \quad (19)$$

and it has already been accounted for in Eq. (14). The term with the fourth power of nucleon distances is

$$P_Q = -\frac{2\pi}{15} m_r^2 \alpha^2 \phi^2(0) \sqrt{\frac{E}{2m_r}} (\vec{\rho}_a - \vec{\rho}_b)^4. \quad (20)$$

The corresponding correction to energy is

$$\begin{aligned} \delta_Q E &= \frac{2\pi}{15} m_r^2 \alpha^2 \phi^2(0) \int_{E_T} dE \sqrt{\frac{E}{2m_r}} \\ &\quad \left[\frac{10}{3} \langle \phi_N | \sum_a \rho_a^2 |E\rangle^2 \right. \\ &\quad - 8 \langle \phi_N | \sum_a \rho_a^i |E\rangle \langle E | \sum_b \rho_b^2 \rho_b^i | \phi_N \rangle \\ &\quad \left. + 4 \langle \phi_N | \left(\sum_a \rho_a^i \rho_a^j - \delta^{ij} \rho_a^2/3 \right) |E\rangle^2 \right] \quad (21) \\ &= \delta_{Q0} E + \delta_{Q1} E + \delta_{Q2} E. \end{aligned}$$

These parts are due to the electric monopole, dipole, and the quadrupole nuclear excitations, correspondingly.

In the second integration region, where $k \sim m$ is large, one performs an expansion in powers of not exactly E , but of the total nuclear energy \tilde{E} ,

$$\tilde{E} = E + \frac{k^2}{2M}, \quad (22)$$

which happens to be much more appropriate. The first expansion term is

$$P = \frac{\pi}{3} m \alpha^2 \phi^2(0) |\vec{r}_a - \vec{r}_b'|^3 \quad (23)$$

and the corresponding correction to energy

$$\delta_Z E = -\frac{\pi}{3} m \alpha^2 \phi^2(0) \sum_{a \neq b}^Z \langle |\vec{r}_a - \vec{r}_b'|^3 \rangle \quad (24)$$

is the modified Zemach moment. The second expansion term for $k \sim m$ is

$$P = \alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 \left(\frac{2m}{k^2} \right)^2 e^{i\vec{k} \cdot (\vec{r}_a - \vec{r}_b')} \tilde{E}. \quad (25)$$

The corresponding nuclear matrix element

$$\langle \phi_N | e^{i\vec{k} \cdot \vec{r}_a} (\tilde{H}_N - E_N) e^{-i\vec{k} \cdot \vec{r}_b} | \phi_N \rangle = \frac{k^2}{2m_N} \delta_{a,b} \quad (26)$$

is proportional to k^2 , so the k integral vanishes after subtraction of singular terms. As a result, no corrections to the modified Zemach moment are found.

Consider now correction due to the finite nucleon size. The proton and neutron charge distribution enters through the convolution with the Coulomb potential in Eq. (5). Since their charge radii are much smaller than that of nuclei, one can perform an expansion of the electric form factors in powers of k^2 . When $k \sim \sqrt{E_m}$, the electric dipole polarizability Eq. (14) is modified by

$$\begin{aligned} P_{\text{FS}} &= \alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 \left(E + \frac{k^2}{2m_r} \right)^{-1} \\ &\quad \times k^2 \frac{(r_{Ea}^2 + r_{Eb}^2)}{6} \frac{[\vec{k}(\vec{\rho}_a - \vec{\rho}_b)]^2}{2} \\ &= -\frac{4\pi}{9} (r_{Ea}^2 + r_{Eb}^2) m_r^2 \alpha^2 \phi^2(0) \sqrt{\frac{E}{2m_r}} (\vec{\rho}_a - \vec{\rho}_b)^2. \end{aligned} \quad (27)$$

The corresponding correction to energy is

$$\begin{aligned} \delta_{\text{FS}} E &= -\frac{16\pi\alpha^2}{9} \phi^2(0) m_r^2 \\ &\quad \times \int_{E_T} dE \sqrt{\frac{E}{2m_r}} \langle \phi_N | \vec{d} | E \rangle \langle E | \vec{\delta} | \phi_N \rangle, \end{aligned} \quad (28)$$

where $\vec{\delta} = \sum_a r_{Ea}^2 \vec{\rho}_a$. When $k \sim m$ the Zemach contribution is corrected by

$$\begin{aligned} \delta_{\text{FZ}} E &= -\frac{\pi}{3} m \alpha^2 \phi^2(0) \sum_a \sum_b^Z \frac{r_{Ea}^2}{3} \langle \vec{\nabla}_a^2 | \vec{r}_a - \vec{r}_b |^3 \rangle \\ &= -\frac{4\pi}{3} m \alpha^2 \phi^2(0) \sum_a \sum_b^Z r_{Ea}^2 \langle |\vec{r}_a - \vec{r}_b| \rangle. \end{aligned} \quad (29)$$

The case $a = b$ is considered separately as it involves a momentum exchange, which is of the order of the inverse of the proton size. When a large momentum is exchanged, the nucleon binding energy can be neglected and the muon sees free nucleons. The individual Zemach radii and nucleon polarizabilities are combined together into effective Dirac-delta type interactions and are accounted for in $\delta_{N,P} E$ in Eqs. (46), and (47).

Consider now corrections from the two-Coulomb exchange using the relativistic (Dirac) Hamiltonian for the muon. Equation (18) is replaced by

$$\begin{aligned} P &= \alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 e^{i\vec{k} \cdot (\vec{r}_a - \vec{r}_b')} \\ &\quad \left(\frac{E_k + m}{2E_k} \frac{1}{\tilde{E} + E_k - m} + \frac{m - E_k}{2E_k} \frac{1}{\tilde{E} + E_k + m} \right), \end{aligned} \quad (30)$$

where $E_k = \sqrt{k^2 + m^2}$. When $k \sim \sqrt{2mE}$ one employs a small k expansion. The leading term coincides with Eq. (18). The next term is

$$P = -\alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 e^{i\vec{k} \cdot (\vec{r}_a - \vec{r}_b')} \frac{\tilde{E} k^2}{(2m\tilde{E} + k^2)^2} \quad (31)$$

Only the quadratic term in nuclear distances contributes, and after subtraction of large k singularities the corresponding correction to energy

$$\begin{aligned} \delta_R E &= \frac{2\pi\alpha^2}{3} \phi^2(0) \int_{E_T} dE \sqrt{\frac{E}{2m_r}} |\langle \phi_N | \vec{d} | E \rangle|^2 \\ &\quad \times \left(1 - 5 \frac{m}{M} \right) + O\left(\frac{m}{M}\right)^2 \end{aligned} \quad (32)$$

is in agreement with the former result of Ref. [7] in the infinite nuclear mass limit.

When $k \sim m$ one can perform the Taylor expansion of the integrand of Eq. (30) in powers of \tilde{E} . The term without \tilde{E} ,

$$P = \alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 e^{i\vec{k} \cdot (\vec{r}_a - \vec{r}_b')} \frac{2m}{k^2}, \quad (33)$$

is exactly the same as in the nonrelativistic limit, and has already been accounted for. The linear in \tilde{E} term in Eq. (30) is

$$P = -\alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 e^{i \vec{k} \cdot (\vec{r}_a - \vec{r}_b)} \tilde{E} \times \frac{m(4m^2 + 3k^2)}{E_k k^4}. \quad (34)$$

The corresponding nuclear matrix element can be transformed using Eq. (26), so correction to energy becomes

$$\delta'_C E = \sum_a \alpha^2 \phi^2(0) \int \frac{d^3 k}{(2\pi)^3} \left(\frac{4\pi}{k^2} \right)^2 \frac{m(4m^2 + 3k^2)}{2m_N E_k k^2} \quad (35)$$

a recoil correction for each individual nucleon. So, for $k \sim m$ the muon sees individual nucleons and these corrections become the sum of nucleon recoil corrections

$$\delta'_C E = -\frac{4}{3m} \alpha^2 \phi^2(0) \left(\frac{Z}{m_N} - \frac{Z^2}{M} \right) \quad (36)$$

with subtracted muon-nucleus recoil correction to avoid the double counting with the so-called pure recoil correction. This is because recoil corrections are by definition included in the Lamb shift as a QED correction for a point nucleus.

The Coulomb exchange is not a complete correction; there are single and double transverse photon exchange corrections, and their calculation is more complicated. The main reason for this is the overlap of the nuclear recoil and nuclear polarizability corrections. Let us repeat now the calculation by replacing the two-Coulomb exchange amplitude, Eq. (30), by a complete two-photon exchange.

When $k \sim E$ or $k \sim \sqrt{E}m$ one can use a dipole approximation, where the coupling of the nucleus to the electromagnetic field is $-\vec{d} \cdot \vec{E}(\vec{R})$, as in Eq. (6). Correction to energy due to two-photon exchange in the dipole approximation is [13]

$$\delta E = -e^4 \phi^2(0) \frac{1}{3} \int_{E_T} dE \langle \phi_D | \vec{d} | E \rangle^2 \int \frac{d\omega}{2\pi i} \int \frac{d^3 k}{(2\pi)^3} \frac{1}{E + \omega + k^2/(2M)} \left(1 + \frac{2\omega^4}{(\omega^2 - k^2)^2} \right) \times \frac{4}{(\omega^2 + 2m\omega - k^2)(\omega^2 - 2m\omega - k^2)}. \quad (37)$$

The leading nonrelativistic term agrees with that in Eq. (14), while the leading relativistic correction agrees with Eq. (32). The higher order correction (in powers of E/m) is

$$\delta'_R E = \frac{4}{3} \alpha^2 \phi^2(0) \langle \phi_N | \vec{d} | \frac{(H_N - E_N)}{m} \times \left[1 + \ln \frac{(H_N - E_N)}{\epsilon} \right] \vec{d} | \phi_N \rangle. \quad (38)$$

Using the commutation relations of Eq. (2) the nuclear matrix element is

$$\frac{4}{3} \langle \phi_N | \vec{d} | (H_N - E_N) \vec{d} | \phi_N \rangle = 2 \left(\frac{Z}{m_N} - \frac{Z^2}{M} \right), \quad (39)$$

and this correction can be rewritten in the form

$$\delta'_R E = \frac{2}{m} \alpha^2 \phi^2(0) \left(1 + \ln \frac{2\bar{E}}{m} + \ln \frac{m}{2\epsilon} \right) \left(\frac{Z}{m_N} - \frac{Z^2}{M} \right), \quad (40)$$

where

$$\ln \bar{E} = \frac{\langle \phi_N | \vec{d} | (H_N - E_N) \ln(H_N - E_N) \vec{d} | \phi_N \rangle}{\langle \phi_N | \vec{d} | (H_N - E_N) \vec{d} | \phi_N \rangle}. \quad (41)$$

The identity (39) is an approximate one. The electric dipole operator does not commute with the nuclear potential. Neglected terms can be interpreted as if due to \vec{A}^2 vertex with pions.

When $k \sim m$ the complete two-photon exchange is a recoil correction from individual protons [see Eq. (32) of Ref. [14]],

$$\delta_H E = \frac{4\pi^2 \alpha^2}{m m_N} \phi^2(0) \int_{\epsilon} \frac{d^3 k}{(2\pi)^3} \times \left[\frac{k^4 + 6k^2 m^2 + 8m^4}{k^6 \sqrt{k^2 + m^2}} - \frac{1}{k^3} \right]. \quad (42)$$

The contribution beyond the previously considered Coulomb part, Eq. (36), is

$$\delta'_H E = \frac{2}{m} \alpha^2 \phi^2(0) \ln \frac{2\epsilon}{m} \left(\frac{Z}{m_N} - \frac{Z^2}{M} \right), \quad (43)$$

where we again subtract the corresponding nuclear recoil correction. The $\ln \epsilon$ dependence cancels out with that in Eq. (40), as it should, and the sum of higher order corrections is

$$\delta_{HO} E = \delta'_C E + \delta'_R E + \delta'_H E = \frac{2}{m} \alpha^2 \phi^2(0) \left(\frac{1}{3} + \ln \frac{2\bar{E}}{m} \right) \left(\frac{Z}{m_N} - \frac{Z^2}{M} \right). \quad (44)$$

These are all nuclear structure corrections up to the order $\alpha^5 m^2/M$. In some cases, such as for the deuteron nucleus, where the magnetic moment is relatively large, higher order effects due to the second-order magnetic interaction $\alpha^5 m^3/M^2$ may play a role. The corresponding correction for the deuteron was obtained in Ref. [7],

$$\delta_M E = \frac{8\pi \alpha^2}{3} \phi^2(0) \left(\frac{g_p - g_n}{4m_p} \right)^2 \times \int_{E_T} dE \sqrt{\frac{E}{2m_r}} \langle \phi_N | \vec{s}_p - \vec{s}_n | E \rangle^2, \quad (45)$$

but the numerical value, as was pointed out in [11], was in error, so we correct it here and present our updated value in Table I.

There are, in addition, contributions due to intrinsic elastic and inelastic two-photon exchanges with individual nucleons, which include the third Zemach moment and the nucleon polarizability. While the contribution from the proton is well known from studies on muonic hydrogen [15], $\Delta E(2S) = -36.9(2.4) \mu\text{eV}$, less is known about the contribution from the neutron. Following [9], we assume that this contribution is as large as the inelastic part for the proton $13.5 \mu\text{eV}$, and associate 50% uncertainty. Therefore, the contribution from intrinsic nucleon polarizabilities and elastic two-photon exchanges is

$$\delta_P E = -\frac{8 Z^4}{n^3} \delta_{l0} \frac{m_{rN}^3}{m_{rH}^3} 36.9 \text{ meV}, \quad (46)$$

$$\delta_N E = -\frac{8(A-Z) Z^3}{n^3} \delta_{l0} \frac{m_{rN}^3}{m_{rH}^3} 13.5 \text{ meV}. \quad (47)$$

The final expression for the nuclear polarizability combined with the elastic contribution but with subtracted nuclear recoil of order $\alpha^5 m^2/M$ is

$$\begin{aligned} \Delta E = & \delta_0 E + \delta_C E + \delta_R E + \delta_Q E + \delta_{FS} E + \delta_{FZ} E \\ & + \delta_M E + \delta_P E + \delta_N E + \delta_{HO} E + \delta_Z E. \end{aligned} \quad (48)$$

and the elastic contribution for the neutron using Galster parametrization [16] is found to be negligible.

Numerical results for muonic deuterium are obtained by using the AV18 potential [17] with the help of a discrete variable representation [18] method for solving the Schrödinger equation, and are presented in Table I. They

TABLE I: Nuclear structure corrections in muonic deuterium for 2P-2S transition. Fundamental physical constants are from Ref. [19], and $r_p^2 = 0.8409^2 \text{ fm}^2$, $r_n^2 = -0.1161 \text{ fm}^2$. $\delta_C^{(0)}$ from [11] includes only the logarithmic part of $\delta_{C1} E$, which we find here to be a good approximation.

Correction	Value in meV	Eq.	Ref. [11]	[11]-AV18
$\delta_0 E$	1.910	(14)	$\delta_{D1}^{(0)}$	1.907
$\delta_{C1} E$	-0.255	(15)	$\rightarrow \delta_C^{(0)}$	-0.262
$\delta_{C2} E$	-0.006	(16)	$\rightarrow \delta_C^{(0)}$	
$\delta_{Q0} E$	-0.042	(21)	$\delta_{R2}^{(2)}$	-0.042
$\delta_{Q1} E$	0.139	(21)	$\delta_{D1D3}^{(2)}$	0.139
$\delta_{Q2} E$	-0.061	(21)	$\delta_Q^{(2)}$	-0.061
$\delta_{FS} E$	0.020	(28)	$\delta_{NS}^{(2)}$	0.015
$\delta_{FZ} E$	-0.018	(29)	$\delta_{np}^{(1)}$	-0.017
$\delta_R E$	-0.026	(32)	$\rightarrow \delta_L^{(0)} + \delta_T^{(0)}$	-0.017
$\delta_{HO} E$	0.004	(44)	$\rightarrow \delta_L^{(0)} + \delta_T^{(0)}$	
$\delta_M E$	-0.008	(45)	$\delta_M^{(0)}$	-0.008
$\delta_P E$	0.043(3)	(46)		0.0135
$\delta_N E$	0.016(8)	(47)		0.0135
ΔE	1.717(20)			1.681(20)

are generally in agreement with our previous calculations

[7] with few exceptions. Differences are due to improved mass dependence of corrections beyond the nonrelativistic dipole term. We also corrected the magnetic contribution, which previously was in error, and included higher order correction $\delta_{HO} E$, and most importantly the polarizability of the neutron. We have also included, following Refs. [10, 11] finite size corrections δ_{FS} and δ_{FZ} , although our results are slightly different.

In comparison to Refs [10, 11], we agree with their numerics, agree with the use of reduced mass in $\delta_Q E$, but disagree with their mass dependence of all other higher order corrections. Moreover, our formula for the finite size correction δ_{FS} slightly disagree with the corresponding $\delta_{NS}^{(2)}$ due to the opposite sign for the neutron radius contribution. Our $\delta_R E$ differs from the corresponding $\delta_{LT}^{(0)}$, apart from mass dependence, due to the fact that for the large momentum exchange $k \sim m$, the dipole approximation does not hold and we account for this in $\delta_{HO} E$. Finally, we consider the elastic contribution of the proton structure correction to be a part of the overall nuclear structure correction, in contrast to Refs. [9, 11]. Our argument is the following. The nuclear structure contribution to high extent is given by the forward two-photon scattering of the nucleus. When momentum exchange is much larger than the nuclear binding energy, muon sees individual nucleons and the total scattering amplitude is a coherent sum of total scattering amplitudes from each nucleon. By total we mean the elastic and inelastic contributions, since this division is a pure convention. Therefore, both contributions should be included in the calculation of the Lamb shift, and we include them, for convenience, in the part called the nuclear structure correction.

Considering the uncertainty related to numerical evaluation of matrix elements, Ref. [11] has performed calculations with AV18 potentials and with various orders of chiral effective field theory, finding 0.6% dependence on the potential used. Our numerical values, when the same formulas are used, are in perfect agreement with those of Ref. [11]. Since we neglect possible corrections to the electric dipole moment, which in fact depends on the model potential, we do not associate additional uncertainty beyond that assumed for the electric dipole polarizability. Regarding uncalculated higher order terms, the most significant seems to be the Coulomb correction to $\delta_Q E$ in Eq. (21), which we estimate by about 0.005 meV. Therefore, the final uncertainty is determined by 50% of the estimated neutron polarizability, 1% of ΔE due to neglect of corrections to the electric dipole moment, and 0.005 meV due to neglected higher order terms.

Our final result for the nuclear structure correction ΔE is not in perfect agreement with that of Ref. [11], as explained above, mostly due to inclusion of the proton elastic contribution. In spite of other small discrepancies with [11], the presented perturbative approach seems to be more efficient than the dispersion relation approach of Ref. [9]. If further improvements are required, the best way is probably by joining within the dispersion

relation approach, the inelastic scattering data at high energies with nuclear model calculations at low energies, to account properly for the high energy structure of the deuteron.

No. 2012/04/A/ST2/00105.

Acknowledgments

The authors would like to acknowledge the support of the Polish National Science Center (NCN) under Grant

[1] R. Pohl *et al.*, *Nature* (London) **466**, 213 (2010).
 [2] A. Antognini *et al.*, *Science* **339**, 417 (2013).
 [3] R. Pohl, R. Gilman, G.A. Miller, and K. Pachucki, *Annu. Rev. Nucl. Part. Sci.* **63**, 175 (2013).
 [4] R. Pohl (private communication).
 [5] J.L. Friar, *Phys. Rev. C* **16**, 1540 (1977).
 [6] W. Leidemann and R. Rosenfelder, *Phys. Rev. C* **51**, 427 (1995).
 [7] K. Pachucki, *Phys. Rev. Lett.* **106**, 193007 (2011).
 [8] J.L. Friar, *Phys. Rev. C* **88**, 034003 (2013).
 [9] C. E. Carlson, M. Gorchtein, and M. Vanderhaeghen, *Phys. Rev. A* **89**, 022504 (2014).
 [10] C. Ji, N. Nevo Dinur, S. Bacca, and N. Barnea, *Phys. Rev. Lett.* **111**, 143402 (2013); N. Nevo Dinur, N. Barnea, C. Ji, and S. Bacca, *Phys. Rev. C* **89**, 064317 (2014); C. Ji, N. Nevo Dinur, S. Bacca, and N. Barnea, *Few-Body Syst.* **55**, 917 (2014).
 [11] O.J. Hernandez, C. Ji, S. Bacca, N. Nevo Dinur, and N. Barnea, *Phys. Lett. B* **736**, 344 (2014).
 [12] A. Wienczek, M. Puchalski, and K. Pachucki, *Phys. Rev. A* **90**, 022508 (2014).
 [13] J. Bernabéu and T.E.O. Ericson, *Z. Phys. A* **309**, 213 (1983).
 [14] K. Pachucki, *J. Phys. B* **31**, 5123 (1998).
 [15] C. E. Carlson and M. Vanderhaeghen, *Phys. Rev. A* **84**, 020102(R) (2011).
 [16] T.R. Gentile and C.B. Crawford, *Phys. Rev. C* **83**, 055203 (2011).
 [17] R.B. Wiringa, V.G.J. Stoks, and R. Schiavilla, *Phys. Rev. C* **51**, 38 (1995).
 [18] D. T. Colbert and W. H. Miller, *J. Chem. Phys.* **96**, 1982 (1992).
 [19] Peter J. Mohr, Barry N. Taylor, and David B. Newell, *Rev. Mod. Phys.* **84**, 1527 (2012).